# Timescale of CH4-CO

## conversion matters

### CH<sub>4</sub>-CO conversion pathways in hot Jupiter atmospheres

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#### What motivated us?

- CH<sub>4</sub> and CO major C-bearing chemical species in hot Jupiter atmospheres.
  CH<sub>4</sub> can be kinetically converted to CO and back.
- There are many  $CH_4$ -CO conversion pathways.
- The pathway that dominates at pressures where  $CH_4$  and CO are interconverted slower than transported by the wind is critical for **quenching**.
- What are the main  $CH_4$ -CO conversion pathways on hot Jupiters?
- Do these pathways change around the planet?

#### How did we model this?

- Unified Model, Met Office's 3D general circulation model (GCM)
- Fluid dynamics coupled with correlated-k radiative transfer and chemical kinetics
  Solar metallicity, aerosol-free
- Planets: HAT-P-11b, HD 189733b, HD 209458b, WASP-17b

#### What did we find?

- The same pathway operates at the  $CH_4$  quench level.
- Different pathways at lower pressures.





#### Transport vs chemistry

- Left: **transport timescale** from our GCM.
- Right: CH<sub>4</sub>-CO conversion timescale from Tsai+2022, A&A (for illustration only).



CH <sub>2</sub> +H <sub>2</sub> O
H <sub>2</sub> +M
H,CO+CO
CO+H <sub>2</sub>
+CO <sup>-</sup>

 $\begin{array}{l} \mathsf{OH+CH}_3(\mathsf{+M}) \rightarrow \mathsf{CH}_3\mathsf{OH}(\mathsf{+M}) \\ \mathsf{CH}_2\mathsf{OH+H}_2 \leftarrow \mathsf{CH}_3\mathsf{OH+H} \\ \mathsf{CH}_2\mathsf{OH+H} \rightarrow \mathsf{H}_2\mathsf{CO+H}_2 \\ \mathsf{H}_2\mathsf{CO+H} \rightarrow \mathsf{HCO+H}_2 \\ \mathsf{HCO+M} \rightarrow \mathsf{H+CO+M} \end{array}$ 

#### How significant is this?

•Using CH<sub>4</sub>-CO chemical relaxation schemes is OK if no variation in CH<sub>4</sub>-CO conversion at quench level is expected.

• But: which CH<sub>4</sub>-CO pathway to use is important as it determines CH<sub>4</sub> and CO abundances at observable pressures.

